

# EPFL

MICRO-517

# Optical Design with ZEMAX OpticStudio

Lecture 3

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### Outline

#### **Theory**

- Process of optical design
- Paraxial model of typical optical instruments
- Paraxial lens design layout and ray tracing

#### **ZEMAX Practice**

- Paraxial design layout and ray tracing
- Functional verification



## The Optical Design Process



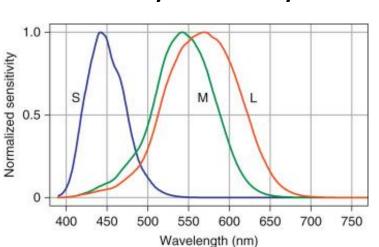
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# **Specification**



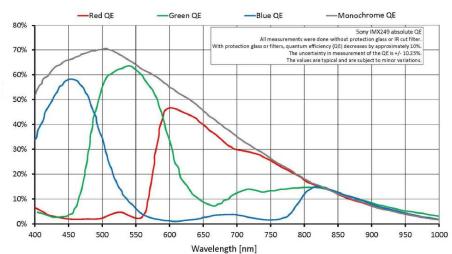
#### **Spectral Transmission Requirement**

Typical normalized spectral sensitivity of human eye

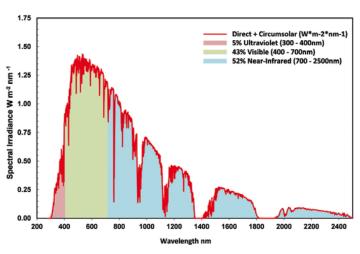


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Typical spectral sensitivity of color image sensor



#### Solar spectral irradiance





## Specification



#### **Imaging**

#### General

- Focal or afocal design
- Type and size of the object
- Spectrum and illuminance (irradiance) of the light
- Distance of the object
- Size and sensitivity of the detector

#### Image quality

- Resolution (periods per mm) as a function of field radius and defocus
- Contrast in relation to detector noise
- Depth of field
- Tolerances on field curvature and distortion
- Tolerances on illumination inhomogeneity
- Telecentricity



## Specification



#### **Paraxial Parameters**

#### Microscope objective

- Field diameter
- Aperture angle or numerical aperture
- Magnification
- Working distance (lens-object)
- Tube length (lens-image distance)
- Length and diameter of the objective

#### **Camera and telescope**

- Focal length
- Back focal length (working distance)
- F-number
- Field angle



## Specification

**Specification** 

Layout

Thin-lens Predesign

Surface Model

**Optimization** 

**Tolerancing** 

#### **Environmental and Economical**

- Physical: temperature, pressure, humidity, shocks, vibrations, dust, radiation, ...
- Chemical: reactants, corrosion, solvents, ...
- Biological: fungi, material toxicity, ...
- Economical: cost, market price, batch size, delivery date, packaging, transport, ...
- Limits in size and weight (potable and space applications)





## Layout

Specification Layout Thin-lens Predesign Surface Model Optimization Tolerancing

- Paraxial phase of a design
- Quicker raytracing computation
- Functional verification
- Corrects primary chromatic aberrations
- Third-order aberrations left out except field curvature
- Complex system can involve lots of algebra



# **Layout**

**Specification** 

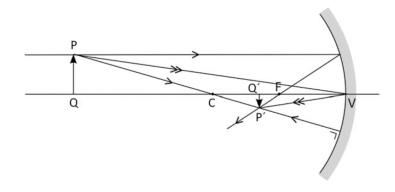
Layout

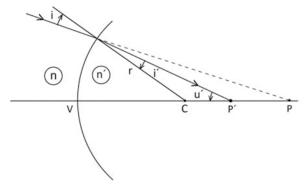
**Thin-lens Predesign**  **Surface** Model

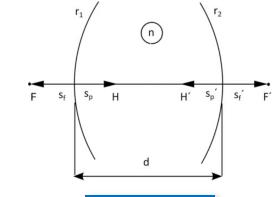
**Optimization** 

**Tolerancing** 

#### **Image Equations Review**







#### **Spherical Mirror**

$$\frac{1}{s'} + \frac{1}{s} = -K \qquad K = -\frac{2}{r} \qquad \qquad \frac{n'}{s'} - \frac{n}{s} = K \qquad K = \frac{n' - n}{r}$$

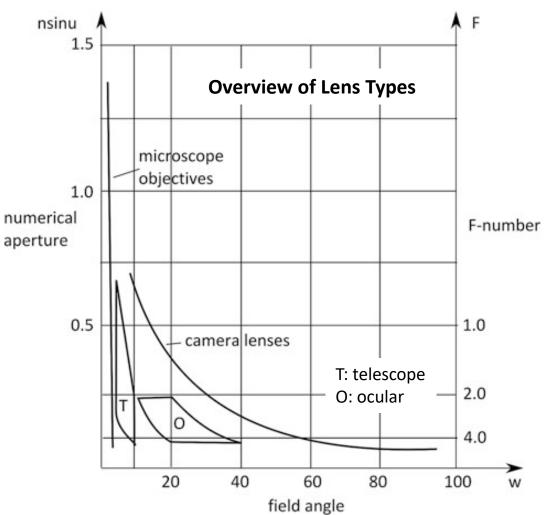
#### **Spherical Refractive Surface**

$$\frac{n'}{s'} - \frac{n}{s} = K$$
  $K = \frac{n' - r}{r}$ 

#### **Single Lens**

$$\frac{1}{s'} - \frac{1}{s} = K$$

$$K = (n-1) \left[ \frac{1}{r_1} - \frac{1}{r_2} + \frac{(n-1)d}{nr_1r_2} \right]$$



#### **Scale factor**

Focal length acts as the scale factor for the system to be designed

$$K = \frac{1}{f} = \frac{1}{h_1} \sum_{i=1}^{k} h_i K_i$$

#### When a system is scaled by a factor p:

- Focal length:  $f \rightarrow pf$
- Radii:  $r_i \rightarrow pr_i$
- Distances:  $d_i \rightarrow pd_i$
- Aperture angles:  $u_i \rightarrow u_i$
- Field angles:  $W_i \rightarrow W_i$

#### Field curvature (Petzval sum)

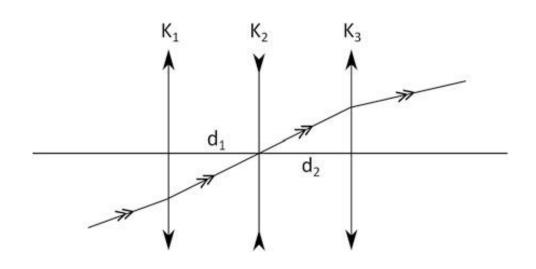
$$P = \sum_{i=1}^{k} \frac{K_i}{n_i}$$

Telescope, microscope: P < K

 $P = \sum_{i=1}^{k} \frac{K_i}{n_i}$  Photographic lens:  $P < 0.2K \sim 0.4K$ 

Lithography objective: no tolerance





- Height of marginal ray at surface i
- Height of chief ray at surface i
- Abbe number at surface i

$$V = \frac{n_d - 1}{\left(n_F - n_C\right)} \quad \begin{array}{ll} n_d & \text{Helium d-line (588 nm)} \\ n_F & \text{Hydrogen F-line (486 nm)} \\ n_C & \text{Hydrogen C-line (656 nm)} \end{array}$$

#### Longitudinal chromatic aberration (LCA) **Transverse chromatic aberration (TCA)**

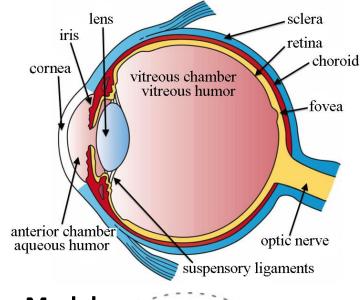
$$C_1 = \sum_{i=1}^{k} h_i^2 K_i / V_i = h_1^2 K / V_s$$

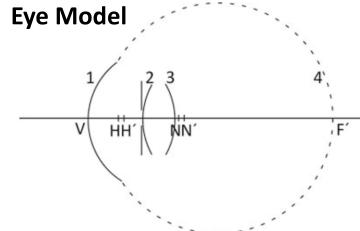
$$C_1 = \sum_{i=1}^k h_i^2 K_i / V_i = h_1^2 \ K / V_s$$
 Equivalent Abbe number 
$$\frac{K_s}{V_s} = \sum_{i=1}^k \frac{K_i}{V_i} \quad K_s = \sum_{i=1}^k K_i$$

$$C_2 = \sum_{i=1}^k h_i \overline{h}_i K_i / V_i$$

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## Typical Optical Instruments | Human Eye



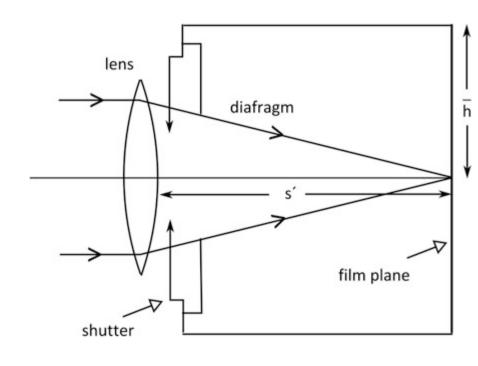


- Cornea:  $\sim$ 50 µm thick, approximately spherical, curvature  $r \sim$  7.8 mm (average)
- Aqueous humor: n ~ 1.336
- Crystalline Lens: ~7 mm behind the cornea, n ~ 1.386 (outside) to 1.406 (center), thickness and curvature controlled by ciliary muscle, power ~17 to 25 diopters
- Iris diaphragm: diameter ~2 to 8 mm
- Vitreous humor: n ~ 1.336.
- Retina: ~25 mm behind the cornea, curvature r ~ 12 mm
- Resolves ~0.1 mm at a distance of 250 mm

	#	Radius	Distance	Index
ОВЈ	0	_	Ş	1
	1	7.8	7	1.336
	2	10.0	3.6	1.413
	3	-6.0	13.4	1.336
IMA	4	-12.0	_	_

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## Typical Optical Instruments | Camera



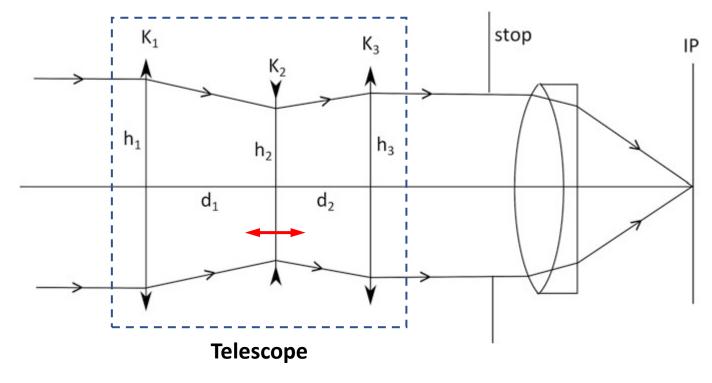
$$\frac{1}{s'} = \frac{1}{s} + \frac{1}{f}$$

F-number F # = f'/DDepth of field (focus)  $\Delta z = \pm \Delta y \cdot F \#$ 

- Huge variety of lens designs
- 30 period/mm standard resolution for modern camera lens

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# Typical Optical Instruments | Zoom Lens



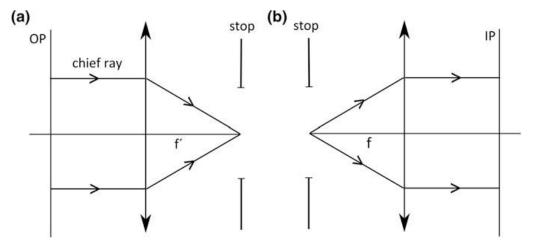
For telescope  $h_1K_1 + h_2K_2 + h_3K_3 = 0$   $u_1 = 0$   $u_2 = -h_1K_1$   $u_3 = h_3K_3$  $h_2 = h_1(1 - K_1d_1) = h_3(1 - K_3d_2)$ 

#### **Angular magnification**

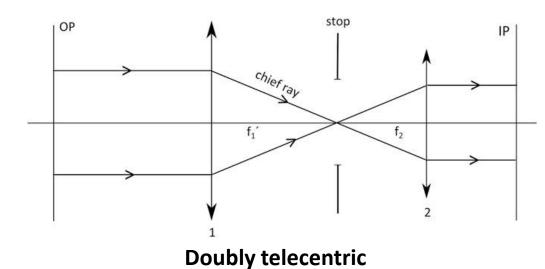
$$M = \frac{h_1}{h_3} = \frac{-K_3}{K_1 + K_2 - \frac{d_1}{d_1} K_1 K_2}$$

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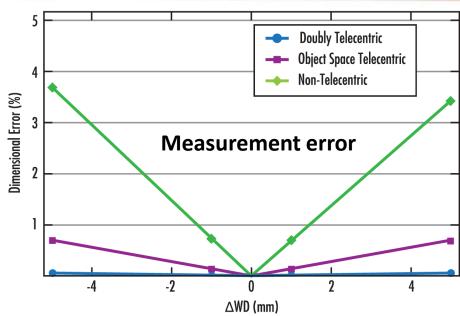
## Typical Optical Instruments | Telecentric Lens



**Object space telecentric** Image space telecentric

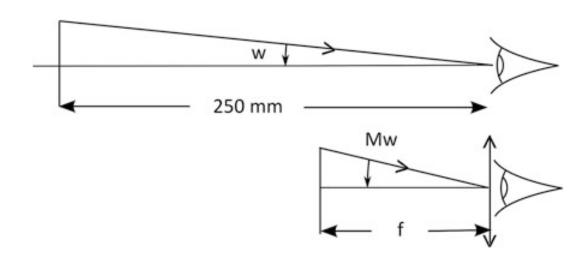








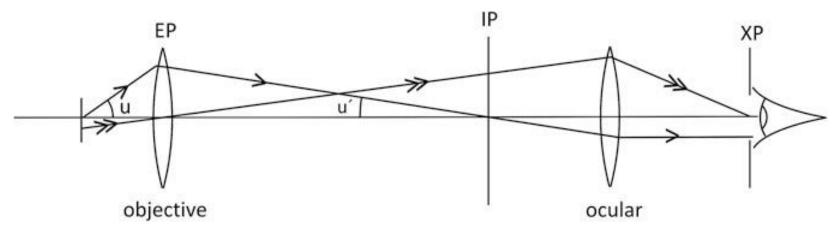
# Typical Optical Instruments | Magnifier



Angular magnification 
$$M = \frac{250}{f}$$



## Typical Optical Instruments | Microscope



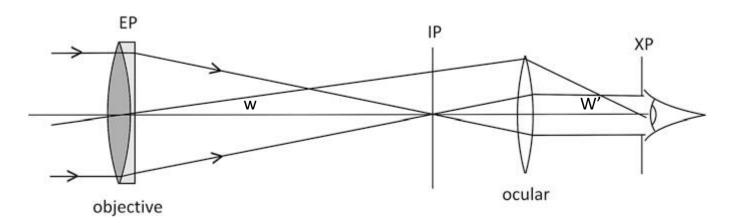
Angular magnification 
$$M=M_{ob} \frac{250}{f_{oc}}$$
 where  $M_{ob}$  is linear magnification of objective

- Objective design is the key
- "Infinity-conjugated" objective
  - Objective plane at front focal plane
  - Require a tube lens for image formation



## Typical Optical Instruments | Telescope

#### **Keplerian telescope**



Linear magnification  $\,M_L = -\,f_2/f_1\,$  Angular magnification  $\,M_A = -\,f_1/f_2\,$ 

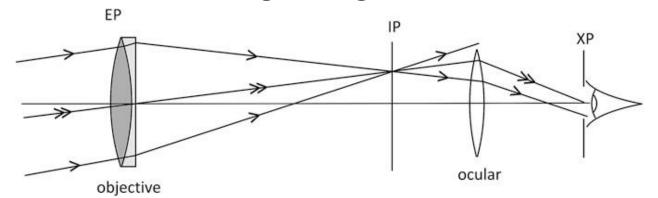
Lagrangian invariant

$$M_A = 1/M_L$$

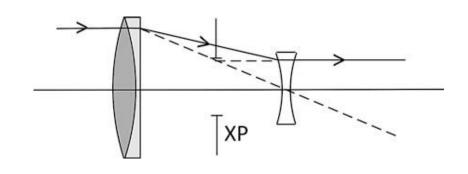


## Typical Optical Instruments | Telescope

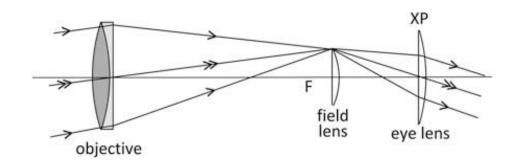
#### **Vignetting**



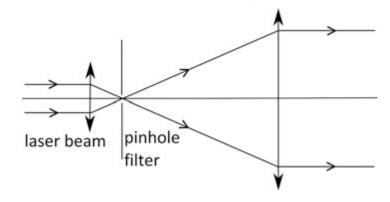
#### Galilei's telescope



#### Vignette correction with field lens



#### Laser beam expander









### Homework

#### **Objective**

- Practice ZEMAX ray tracing with paraxial models
- Familiarize yourself with principles of various optical instruments
- Understand functional verification of your design
- 1. Design a paraxial telescope of magnification M = 100 and objective lens aperture diameter of 300 mm. Note that Zemax OpticStudio always works with semi-diameter. Use an appropriate field lens to improve vignetting.
- 2. Design a paraxial microscope of magnification M = 100. The microscope objective works at "infinite-conjugate" (object at the front focal plane). Use a tube lens of 200 mm.
- 3. Design a paraxial doubly telecentric lens with magnification M = 0.1. Use field heights of 0, 30, 100 mm at an object distance of 200 mm.
- 4. Design a paraxial zoom lens with the range of angular magnification M = 2 5. Assume object at infinity. Use field angle of 0, 3, and 5 degrees. Use another paraxial lens as the last focusing lens.

You should submit the ZEMAX design files (.zos). No need to generate spot diagram since we are in a paraxial regime. However, it may help you diagnose design problems at various stages.